



Existence of a class of rotopulsators



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ABSTRACT

We prove the existence of a class of rotopulsators for the n -body problem in spaces of constant Gaussian curvature of dimension $k \geq 2$.

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1. Introduction

By n -body problems, we mean problems where we want to find the dynamics of n point particles. If the space in which such a problem is defined is a space of zero Gaussian curvature, then we call any solution to such a problem for which the point particles describe the vertices of a polytope that retains its shape over time (but not necessarily its size) a homographic orbit.

A rotopulsator, also known as a rotopulsating orbit, is a type of solution to an n -body problem for spaces of constant Gaussian curvature $\kappa \neq 0$ that extends the definition of homographic orbits to spaces of constant Gaussian curvature (see [7]).

Homographic orbits (and therefore rotopulsators) can be used to determine the geometry of the universe locally (see for example [4,7]).

In this paper, we will prove the existence of a subclass of rotopulsators that form a natural generalization of orbits found in [4,6].

While the orbits in [4,6] are referred to as homographic, it has been argued in [7] that using the term ‘homographic orbit’ for spaces of constant Gaussian curvature makes little sense. We will therefore, following [7], speak of rotopulsating orbits instead.

While this paper mainly builds on results obtained in [4,6,22], research on n -body problems for spaces of constant Gaussian curvature goes back to Bolyai [1] and Lobachevsky [19], who independently proposed a curved 2-body problem in hyperbolic space \mathbb{H}^3 in the 1830s. In later years, n -body problems for spaces of constant Gaussian curvature have been studied by mathematicians such as Dirichlet, Schering [20,21], Killing [12–14] and Liebmann [16–18]. More recent results were obtained by Kozlov and Harin [15], but the study of n -body problems in spaces of constant Gaussian curvature for the case that $n \geq 2$ started with [9–11] by Diacu, Pérez-Chavela, and Santoprete. Further results for the $n \geq 2$ case were then

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obtained by Cariñena, Rañada, and Santander [2], Diacu [3,4,6], Diacu and Kordlou [7], and Diacu and Pérez-Chavela [8]. For a more detailed historical overview, please see [4,6,5,7], or [9].

In this paper, we will prove the following two theorems:

Theorem 1.1. For any rotopulsating solution of (2.2) formed by vectors $\{\mathbf{q}_i\}_{i=1}^n$ as defined in (2.3), the vectors $\{\mathbf{Q}_i\}_{i=1}^n$ have to form a regular polygon if ρ is non-constant.

Theorem 1.2. Rotopulsating orbits formed by vectors $\{\mathbf{q}_i\}_{i=1}^n$ as defined in (2.3) exist if the vectors $\{\mathbf{Q}_i\}_{i=1}^n$ form a regular polygon.

To prove these theorems, we will use a method strongly inspired by [4,6,22]. Specifically, we will first deduce a necessary and sufficient criterion for the existence of rotopulsators. This will be done in Section 2. We will then prove Theorems 1.1 and 1.2 in Sections 3 and 4 respectively.

2. A criterion for the existence of rotopulsators

In this section, we will formulate a necessary and sufficient criterion for the existence of rotopulsating orbits of the type described in (2.3).

Consider the n -body problem in spaces of constant Gaussian curvature $\kappa \neq 0$.

As has been shown in [5], we may assume that κ equals either -1 , or 1 .

We will denote the masses of its n point particles to be $m_1, m_2, \dots, m_n > 0$ and their positions by the k -dimensional vectors

$$\mathbf{q}_i^T = (q_{i1}, q_{i2}, \dots, q_{ik}) \in \mathbb{M}_\kappa^{k-1}, \quad i = \overline{1, n}$$

where

$$\mathbb{M}_\kappa^{k-1} = \{(x_1, x_2, \dots, x_k) \in \mathbb{R}^k | \kappa(x_1^2 + x_2^2 + \dots + x_{k-1}^2 + \sigma x_k^2) = 1\}, \quad k \in \mathbb{N}$$

and

$$\sigma = \begin{cases} 1 & \text{for } \kappa > 0 \\ -1 & \text{for } \kappa < 0. \end{cases}$$

Furthermore, consider for m -dimensional vectors $\mathbf{a} = (a_1, a_2, \dots, a_m)$, $\mathbf{b} = (b_1, b_2, \dots, b_m)$ the inner product

$$\mathbf{a} \odot_m \mathbf{b} = a_1 b_1 + a_2 b_2 + \dots + a_{m-1} b_{m-1} + \sigma a_m b_m. \tag{2.1}$$

Then, following [3,4,6,9–11] and the assumption that $\kappa = \pm 1$ from [5], we define the equations of motion for the curved n -body problem as the dynamical system described by

$$\ddot{\mathbf{q}}_i = \sum_{j=1, j \neq i}^n \frac{m_j [\mathbf{q}_j - (\sigma \mathbf{q}_i \odot_k \mathbf{q}_j) \mathbf{q}_i]}{[\sigma - (\mathbf{q}_i \odot_k \mathbf{q}_j)^2]^{\frac{3}{2}}} - (\sigma \dot{\mathbf{q}}_i \odot_k \dot{\mathbf{q}}_i) \mathbf{q}_i, \quad i = \overline{1, n}. \tag{2.2}$$

Let

$$T(t) = \begin{pmatrix} \cos(\theta(t)) & -\sin(\theta(t)) \\ \sin(\theta(t)) & \cos(\theta(t)) \end{pmatrix}$$

be a 2×2 rotation matrix, where $\theta(t)$ is some real valued, twice continuously differentiable, scalar function, for which $\theta(0) = 0$.

We will consider rotopulsating orbit solutions of (2.2) of the form

$$\mathbf{q}_i(t) = \begin{pmatrix} \rho(t) T(t) \mathbf{Q}_i \\ Z(t) \end{pmatrix} \tag{2.3}$$

where $\mathbf{Q}_i \in \mathbb{R}^2$ is a constant vector and $Z(t) \in \mathbb{R}^{k-2}$ is a twice differentiable, vector valued function.

Finally, before formulating our criterion, we need to introduce some notation and a lemma:

Let $m \in \mathbb{N}$. Let $\langle \cdot, \cdot \rangle_m$ be the Euclidean inner product on \mathbb{R}^m and let $\|\cdot\|_m$ be the Euclidean norm on \mathbb{R}^m . Let $i, j \in \{1, \dots, n\}$. By construction $\|\mathbf{Q}_i\|_2 = \|\mathbf{Q}_j\|_2$ for all $i, j \in \{1, \dots, n\}$ and we will assume that $\|\mathbf{Q}_i\|_2 = 1$. Let β_i be the angle between \mathbf{Q}_i and the first coordinate axis. The lemma we will need to prove our criterion is:

Lemma 2.1. The functions ρ and θ , are related through the following formula: $\rho^2(t) \dot{\theta}(t) = \rho^2(0) \dot{\theta}(0)$.

Proof. In [6], using the wedge product, Diacu proved that

$$\sum_{i=1}^n m_i \dot{\mathbf{q}}_i \wedge \mathbf{q}_i = \mathbf{c}$$

where \mathbf{c} is a constant bivector.

If $\{\mathbf{e}_i\}_{i=1}^k$ are the standard base vectors in \mathbb{R}^k , then we can write \mathbf{c} as

$$\mathbf{c} = \sum_{i=1}^k \sum_{j=1}^k c_{ij} \mathbf{e}_i \wedge \mathbf{e}_j \tag{2.4}$$

where $\{c_{ij}\}_{i=1, j=1}^k$ are constants. As $\mathbf{e}_i \wedge \mathbf{e}_j = -\mathbf{e}_j \wedge \mathbf{e}_i$ and $\mathbf{e}_i \wedge \mathbf{e}_i = 0$ (see [6]), for $i, j \in \{1, \dots, n\}$, we can rewrite (2.4) as

$$\mathbf{c} = \sum_{i=1}^k \sum_{j=i+1}^k C_{ij} \mathbf{e}_i \wedge \mathbf{e}_j \tag{2.5}$$

where $C_{ij} = c_{ij} - c_{ji}$.

Calculating C_{12} will give us our result:

Note that

$$T^T = T^{-1} \quad \text{and} \quad \dot{T} = \dot{\theta} \begin{pmatrix} 0 & -1 \\ 1 & 0 \end{pmatrix} T \tag{2.6}$$

and

$$\begin{aligned} C_{12} &= \sum_{i=1}^n m_i (q_{i1} \dot{q}_{i2} - q_{i2} \dot{q}_{i1}) \\ &= \sum_{i=1}^n m_i (q_{i1}, q_{i2}) \begin{pmatrix} 0 & 1 \\ -1 & 0 \end{pmatrix} \begin{pmatrix} \dot{q}_{i1} \\ \dot{q}_{i2} \end{pmatrix}. \end{aligned} \tag{2.7}$$

Using (2.3) with (2.7) gives

$$C_{12} = \sum_{i=1}^n m_i \rho^2 (Q_{i1}, Q_{i2}) T^T \begin{pmatrix} 0 & 1 \\ -1 & 0 \end{pmatrix} \dot{T} \begin{pmatrix} Q_{i1} \\ Q_{i2} \end{pmatrix} + \sum_{i=1}^n m_i \rho \dot{\rho} (Q_{i1}, Q_{i2}) T^T \begin{pmatrix} 0 & 1 \\ -1 & 0 \end{pmatrix} T \begin{pmatrix} Q_{i1} \\ Q_{i2} \end{pmatrix}. \tag{2.8}$$

Note that

$$\rho \dot{\rho} (Q_{i1}, Q_{i2}) T^T \begin{pmatrix} 0 & 1 \\ -1 & 0 \end{pmatrix} T \begin{pmatrix} Q_{i1} \\ Q_{i2} \end{pmatrix} = \frac{\dot{\rho}}{\rho} (q_{i1}, q_{i2}) \begin{pmatrix} 0 & 1 \\ -1 & 0 \end{pmatrix} \begin{pmatrix} q_{i1} \\ q_{i2} \end{pmatrix} = 0.$$

So, using (2.6) repeatedly, we get that

$$\begin{aligned} C_{12} &= \sum_{i=1}^n m_i \rho^2 \dot{\theta} (Q_{i1}, Q_{i2}) T^T \begin{pmatrix} 0 & 1 \\ -1 & 0 \end{pmatrix} \begin{pmatrix} 0 & -1 \\ 1 & 0 \end{pmatrix} T \begin{pmatrix} Q_{i1} \\ Q_{i1} \end{pmatrix} + 0 \\ &= \sum_{i=1}^n m_i \rho^2 \dot{\theta} (Q_{i1}, Q_{i2}) T^T T \begin{pmatrix} Q_{i1} \\ Q_{i2} \end{pmatrix} \\ &= \sum_{i=1}^n m_i \rho^2 \dot{\theta} (Q_{i1}, Q_{i2}) \begin{pmatrix} Q_{i1} \\ Q_{i2} \end{pmatrix} \end{aligned}$$

which means that

$$C_{12} = \rho^2 \dot{\theta} \sum_{i=1}^n m_i (Q_{i1}^2 + Q_{i2}^2). \tag{2.9}$$

As, by construction

$$\sum_{i=1}^n m_i (Q_{i1}^2 + Q_{i2}^2) > 0,$$

we may divide both sides of (2.9) by

$$\sum_{i=1}^n m_i (Q_{i1}^2 + Q_{i2}^2),$$

which gives that

$$\rho^2 \dot{\theta} = \frac{C_{12}}{\sum_{i=1}^n m_i (Q_{i1}^2 + Q_{i2}^2)},$$

which is constant, so $\rho^2 \dot{\theta} = \rho^2(0) \dot{\theta}(0)$. \square

We now have the following necessary and sufficient criterion for the existence of a rototulsating orbit, as described in (2.3):

Criterion 1. Let

$$b_i = \sum_{j=1, j \neq i}^n \frac{m_j (1 - \cos(\beta_i - \beta_j))^{-\frac{1}{2}}}{(2 - \sigma \rho^2 (1 - \cos(\beta_i - \beta_j)))^{\frac{3}{2}}}. \tag{2.10}$$

Then necessary and sufficient conditions for the existence of a rototulsating orbit of non-constant size are that $b_1 = b_2 = \dots = b_n$ and

$$0 = \sum_{j=1, j \neq i}^n \frac{m_j \sin(\beta_i - \beta_j)}{(1 - \cos(\beta_i - \beta_j))^{\frac{3}{2}} (2 - \sigma \rho^2 (1 - \cos(\beta_i - \beta_j)))^{\frac{3}{2}}} \tag{2.11}$$

for all $i \in \{1, \dots, n\}$.

Proof. Note that

$$\dot{T} = \dot{\theta} T \begin{pmatrix} 0 & -1 \\ 1 & 0 \end{pmatrix} \tag{2.12}$$

and consequently

$$\ddot{T} = \ddot{\theta} T \begin{pmatrix} 0 & -1 \\ 1 & 0 \end{pmatrix} - \dot{\theta}^2 T. \tag{2.13}$$

Inserting (2.3) into (2.2) and using (2.12) and (2.13) gives for the first and second lines of (2.2) that

$$\begin{aligned} T \left(\ddot{\rho} I_2 + 2\dot{\rho}\dot{\theta} \begin{pmatrix} 0 & -1 \\ 1 & 0 \end{pmatrix} + \rho \left(\ddot{\theta} \begin{pmatrix} 0 & -1 \\ 1 & 0 \end{pmatrix} - \dot{\theta}^2 I_2 \right) \right) \mathbf{Q}_i \\ = \rho T \left(\sum_{j=1, j \neq i}^n \frac{m_j [\mathbf{Q}_j - (\sigma \mathbf{q}_i \odot_k \mathbf{q}_j) \mathbf{Q}_i]}{[\sigma - (\mathbf{q}_i \odot_k \mathbf{q}_j)^2]^{\frac{3}{2}}} - (\sigma \dot{\mathbf{q}}_i \odot_k \dot{\mathbf{q}}_i) \mathbf{Q}_i \right) \end{aligned} \tag{2.14}$$

where I_2 is the 2×2 identity matrix.

For the last $k - 2$ lines, we get

$$\ddot{Z} = \left(\sum_{j=1, j \neq i}^n \frac{m_j [1 - (\sigma \mathbf{q}_i \odot_k \mathbf{q}_j)]}{[\sigma - (\mathbf{q}_i \odot_k \mathbf{q}_j)^2]^{\frac{3}{2}}} - (\sigma \dot{\mathbf{q}}_i \odot_k \dot{\mathbf{q}}_i) \right) Z. \tag{2.15}$$

Note that

$$\mathbf{q}_i \odot_k \mathbf{q}_j = \rho^2 \langle \mathbf{Q}_j, \mathbf{Q}_i \rangle_2 + Z \odot_{k-2} Z. \tag{2.16}$$

As we have that $\langle \mathbf{Q}_j, \mathbf{Q}_i \rangle_2 = 1$ and as by (2.16),

$$\sigma^{-1} = \mathbf{q}_i \odot_k \mathbf{q}_i = \rho^2 \langle \mathbf{Q}_i, \mathbf{Q}_i \rangle_2 + Z \odot_{k-2} Z,$$

we may rewrite (2.16) as

$$\mathbf{q}_i \odot_k \mathbf{q}_j = \sigma^{-1} + \rho^2 \langle \mathbf{Q}_j, \mathbf{Q}_i \rangle_2 - \rho^2,$$

which can, in turn, be written as

$$\mathbf{q}_i \odot_k \mathbf{q}_j = \sigma^{-1} + \rho^2 (\cos(\beta_i - \beta_j) - 1). \tag{2.17}$$

Furthermore,

$$\dot{\mathbf{q}}_i \odot_k \dot{\mathbf{q}}_i = \langle \dot{\rho} T \mathbf{Q}_i + \rho \dot{T} \mathbf{Q}_i, \dot{\rho} T \mathbf{Q}_i + \rho \dot{T} \mathbf{Q}_i \rangle_2 + \dot{Z} \odot_{k-2} \dot{Z}. \tag{2.18}$$

As T is a rotation in \mathbb{R}^2 , it is a unitary map, meaning that for $v, w \in \mathbb{R}^2$, $\langle Tv, Tw \rangle_2 = \langle v, w \rangle_2$, meaning that (2.18) can be written as

$$\dot{\mathbf{q}}_i \circledast_k \dot{\mathbf{q}}_i = \langle \dot{\rho} \mathbf{Q}_i + \rho T^{-1} \dot{T} \mathbf{Q}_i, \dot{\rho} \mathbf{Q}_i + \rho T^{-1} \dot{T} \mathbf{Q}_i \rangle_2 + \dot{Z} \circledast_{k-2} \dot{Z}. \tag{2.19}$$

Using (2.12) with (2.19) gives

$$\begin{aligned} \dot{\mathbf{q}}_i \circledast_k \dot{\mathbf{q}}_i &= \dot{\rho}^2 + 2\rho\dot{\rho}\dot{\theta} \left\langle \mathbf{Q}_i, \begin{pmatrix} 0 & -1 \\ 1 & 0 \end{pmatrix} \mathbf{Q}_i \right\rangle_2 + \rho^2 \dot{\theta}^2 \|\mathbf{Q}_i\|^2 + \dot{Z} \circledast_{k-2} \dot{Z} \\ &= \dot{\rho}^2 + 0 + \rho^2 \dot{\theta}^2 + \dot{Z} \circledast_{k-2} \dot{Z}. \end{aligned} \tag{2.20}$$

Inserting (2.20) and (2.17) into (2.14) and multiplying both sides by T^{-1} provides us with

$$\begin{aligned} &\left((\ddot{\rho} - \rho\dot{\theta}^2)I_2 + (2\dot{\rho}\dot{\theta} + \rho\ddot{\theta}) \begin{pmatrix} 0 & -1 \\ 1 & 0 \end{pmatrix} \right) \mathbf{Q}_i \\ &= \rho \sum_{j=1, j \neq i}^n \frac{m_j [\mathbf{Q}_j - (1 - \sigma\rho^2(1 - \cos(\beta_i - \beta_j)))\mathbf{Q}_i]}{[\rho^2(1 - \cos(\beta_i - \beta_j))(2 - \sigma\rho^2(1 - \cos(\beta_i - \beta_j)))]^{\frac{3}{2}}} \\ &\quad - (\sigma\rho\dot{\rho}^2 + \sigma\rho^3\dot{\theta}^2 + \sigma\rho\dot{Z} \circledast_{k-2} \dot{Z}) \mathbf{Q}_i. \end{aligned} \tag{2.21}$$

Taking the Euclidean inner product with \mathbf{Q}_i on both sides of (2.21) and using that $\|\mathbf{Q}_i\|_2 = \|\mathbf{Q}_j\|_2 = 1$ provides us with

$$\ddot{\rho} - \rho\dot{\theta}^2 + \sigma\rho\dot{\rho}^2 + \sigma\rho^3\dot{\theta}^2 + \sigma\rho\dot{Z} \circledast_{k-2} \dot{Z} = \left(\sigma - \frac{1}{\rho^2} \right) \sum_{j=1, j \neq i}^n \frac{m_j \left[(1 - \cos(\beta_i - \beta_j))^{-\frac{1}{2}} \right]}{[(2 - \sigma\rho^2(1 - \cos(\beta_i - \beta_j)))]^{\frac{3}{2}}}. \tag{2.22}$$

Taking the Euclidean inner product of (2.21) with $\begin{pmatrix} 0 & 1 \\ -1 & 0 \end{pmatrix} \mathbf{Q}_i$ and using that $\|\mathbf{Q}_i\|_2 = \|\mathbf{Q}_j\|_2 = 1$ gives us that

$$2\dot{\rho}\dot{\theta} + \rho\ddot{\theta} = \sum_{j=1, j \neq i}^n \frac{m_j \sin(\beta_i - \beta_j)}{[(1 - \cos(\beta_i - \beta_j))(2 - \sigma\rho^2(1 - \cos(\beta_i - \beta_j)))]^{\frac{3}{2}}}. \tag{2.23}$$

Let

$$\begin{aligned} b_i &:= \sum_{j=1, j \neq i}^n \frac{m_j \left[(1 - \cos(\beta_i - \beta_j))^{-\frac{1}{2}} \right]}{[(2 - \sigma\rho^2(1 - \cos(\beta_i - \beta_j)))]^{\frac{3}{2}}} \quad \text{and} \\ c_i &:= \sum_{j=1, j \neq i}^n \frac{-m_j \sin(\beta_i - \beta_j)}{[(1 - \cos(\beta_i - \beta_j))(2 - \sigma\rho^2(1 - \cos(\beta_i - \beta_j)))]^{\frac{3}{2}}}. \end{aligned}$$

Inserting (2.20) and (2.17) into (2.15), combined with (2.22) and (2.23), gives the following system of differential equations:

$$\begin{cases} \ddot{\rho} = \rho\dot{\theta}^2 - \sigma\rho\dot{\rho}^2 - \sigma\rho^3\dot{\theta}^2 - \sigma\rho\dot{Z} \circledast_{k-2} \dot{Z} + \left(\sigma - \frac{1}{\rho^2} \right) b_i \\ \ddot{\theta} = \frac{c_i}{\rho} - 2\frac{\dot{\rho}}{\rho}\dot{\theta} \\ \ddot{Z} = (b_i - \sigma\dot{\rho}^2 - \sigma\rho^2\dot{\theta}^2 - \sigma\dot{Z} \circledast_{k-2} \dot{Z}) Z. \end{cases} \tag{2.24}$$

For (2.24) to make sense, we need that

$$b_1 = \dots = b_n \quad \text{and} \quad c_1 = \dots = c_n \tag{2.25}$$

which shows the necessity of (2.25).

Furthermore, that (2.24) has a global solution holds by the same argument as the argument used in the proof of Criterion 1 in [4] to prove global existence of a solution of (15) and (17). By the uniqueness of solutions to ordinary differential equations given suitable initial conditions, the solution to (2.24) must be a rotopulsating orbit, as every step from (2.14) and (2.15) to (2.24) is invertible.

Thus (2.25) is both necessary and sufficient. Finally, as by Lemma 2.1 $\rho^2\dot{\theta} = \rho^2(0)\dot{\theta}(0)$, we have that $\frac{d}{dt}(\rho^2\dot{\theta}) = 0$, which means that the left hand side of (2.23) equals zero, which means that $c_i = 0$. This completes the proof. \square

3. Proof of Theorem 1.1

Criterion 1 in [4] tells us that for $n \geq 3$ bodies of masses $m_1, m_2, \dots, m_n > 0$ moving on the surface \mathbf{M}_k^2 , necessary and sufficient conditions for a rotopulsating orbit as described by (2.3) to be a solution of Eq. (2.2) are given by the equations

$$\delta_1 = \delta_2 = \dots = \delta_n \quad \text{and} \quad \gamma_1 = \gamma_2 = \dots = \gamma_n,$$

where

$$\delta_i = \sum_{j=1, j \neq i}^n m_j \mu_{ji}, \quad \gamma_i = \sum_{j=1, j \neq i}^n m_j \nu_{ji}, \quad i = \overline{1, n},$$

$$\mu_{ji} = \frac{1}{c_{ji}^{\frac{1}{2}} (2 - c_{ji} k r^2)^{\frac{3}{2}}}, \quad \nu_{ji} = \frac{s_{ji}}{c_{ji}^{\frac{3}{2}} (2 - c_{ji} k r^2)^{\frac{3}{2}}}$$

$$s_{ji} = \sin(\alpha_j - \alpha_i), \quad c_{ji} = 1 - \cos(\alpha_j - \alpha_i), \quad i, j = \overline{1, n}, \quad i \neq j$$

where $\alpha_i = \beta_i$, $r = \rho$, $\delta_i := b_i$ and $\gamma_i := c_i$. Thus the conditions of Criterion 1 are exactly the conditions of Criterion 1 in [4] with the added bonus that $\gamma_i = 0$. The proof of Theorem 1.1 in [22] is therefore a proof for Theorem 1.1 as well.

Remark 3.1. It should be noted that in [4,22] rotopulsators as described by (2.3) are called ‘polygonal homographic orbits’. However, as in [7] it was argued that the term ‘homographic’ should be replaced by ‘rotopulsating’, as we did so here.

4. Proof of Theorem 1.2

Let again $r := \rho$, $\alpha_i := \beta_i$, $\delta_i := b_i$ and $\gamma_i := c_i$ in Criterion 1. Then, as before in the proof of Theorem 1.1, the conditions of Criterion 1 become exactly the conditions of Criterion 1 in [4] with the added bonus that $\gamma_i = 0$. In [4], Diacu proved that

Theorem 1. Consider the curved n -body problem, $n \geq 3$, given by (2.2). If n bodies of equal masses, $m := m_1 = m_2 = \dots = m_n$, lie initially at the vertices of a regular n -gon parallel with the (x, y) -plane, then there is a class of initial velocities for which the corresponding solutions are rotopulsators. These orbits also satisfy the equalities $\gamma_1 = \gamma_2 = \dots = \gamma_n = 0$.

and

Theorem 2. If the masses m_1, \dots, m_n , $n \geq 3$, form a rotopulsating solution of the curved n -body problem given by Eq. (2.2), such that the polygon is regular, then $m_1 = m_2 = \dots = m_n$.

The proofs in [4] of Theorems 1 and 2 use Criterion 1 of [4] alone, thus as Criterion 1 of [4] and Criterion 1 coincide, Theorem 1.2 now follows directly from Theorems 1 and 2 in [4].

Remark 4.1. It should be noted that the expressions ‘rotopulsators’ and ‘rotopulsating solution’ as described by (2.3) are not used in Theorems 1 and 2 in [4]. In [4], the expressions used are ‘homographic’ and ‘polygonal homographic solution’ respectively. However, as in [7] it was argued that the term ‘homographic’ should be replaced by ‘rotopulsating’, as we did so here.

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References

- [1] W. Bolyai, J. Bolyai, Geometrische Untersuchungen, Hrsg. P. Stäckel, Teubner, Leipzig, Berlin, 1913.
- [2] J.F. Cariñena, M.F. Rañada, M. Santander, Central potentials on spaces of constant curvature: the Kepler problem on the two-dimensional sphere \mathbb{S}^2 and the hyperbolic plane \mathbb{H}^2 , J. Math. Phys. 46 (2005) 052702.
- [3] F. Diacu, On the singularities of the curved n -body problem, Trans. Amer. Math. Soc. 363 (2011) 2249–2264.
- [4] F. Diacu, Polygonal homographic orbits of the curved n -body problem, Trans. Amer. Math. Soc. 364 (5) (2012) 2783–2802.
- [5] F. Diacu, Relative Equilibria of the Curved N -Body Problem, in: Atlantis Studies in Dynamical Systems, vol. 1, Atlantis Press, Amsterdam, 2012.
- [6] F. Diacu, Relative equilibria in the 3-dimensional curved n -body problem, Mem. Amer. Math. Soc. (in press).
- [7] F. Diacu, S. Kordlou, Rotopulsators of the curved N -body problem, p. 40. arXiv:1210.4947v2.
- [8] F. Diacu, E. Pérez-Chavela, Homographic solutions of the curved 3-body problem, J. Differential Equations 250 (2011) 340–366.
- [9] F. Diacu, E. Pérez-Chavela, M. Santoprete, The n -body problem in spaces of constant curvature, p. 54. arXiv:0807.1747.
- [10] F. Diacu, E. Pérez-Chavela, M. Santoprete, The n -body problem in spaces of constant curvature, Part I: relative equilibria, J. Nonlinear Sci. 22 (2) (2012) 247–266. <http://dx.doi.org/10.1007/s00332-011-9116-z>.
- [11] F. Diacu, E. Pérez-Chavela, M. Santoprete, The n -body problem in spaces of constant curvature, Part II: singularities, J. Nonlinear Sci. 22 (2) (2012) 267–275. <http://dx.doi.org/10.1007/s00332-011-9117-y>.
- [12] W. Killing, Die Rechnung in den nichteuklidischen Raumformen, J. Reine Angew. Math. 89 (1880) 265–287.
- [13] W. Killing, Die Mechanik in den nichteuklidischen Raumformen, J. Reine Angew. Math. 98 (1885) 1–48.
- [14] W. Killing, Die Nicht-Euklidischen Raumformen in Analytischer Behandlung, Teubner, Leipzig, 1885.

- [15] V.V. Kozlov, A.O. Harin, Kepler's problem in constant curvature spaces, *Celestial Mech. Dynam. Astronom.* 54 (1992) 393–399.
- [16] H. Liebmann, Die Kegelschnitte und die Planetenbewegung im nichteuklidischen Raum, *Ber. Königl. Sächs. Gesell. Wiss., Math. Phys. Kl.* 54 (1902) 393–423.
- [17] H. Liebmann, Über die Zentralbewegung in der nichteuklidische Geometrie, *Ber. Königl. Sächs. Gesell. Wisse., Math. Phys. Kl.* 55 (1903) 146–153.
- [18] H. Liebmann, *Nichteuklidische Geometrie*, G.J. Göschen, Leipzig, 1905. 2nd ed. 1912; 3rd ed. Walter de Gruyter, Berlin Leipzig, 1923.
- [19] N.I. Lobachevsky, The new foundations of geometry with full theory of parallels, 1835–1838, in: *Collected Works*, Vol. 2, GITTL, Moscow, 1949, p. 159 (in Russian).
- [20] E. Schering, Die Schwerkraft im Gaussischen Räume, *Nachr. Königl. Gesell. Wiss. Göttingen* 15 (1873) 311–321.
- [21] E. Schering, Die Schwerkraft in mehrfach ausgedehnten Gaussischen und Riemmanschen Räumen, *Nachr. Königl. Gesell. Wiss. Göttingen* 6 (1873) 149–159.
- [22] P. Tibboel, Polygonal homographic orbits in spaces of constant curvature, *Proc. Amer. Math. Soc.* 141 (2013) 1465–1471.